

# FFSWK User Manual (version 0.1)

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## 1 Introduction

This package includes programs for calculation of Finite-Frequency Surface Wave Kernels (FFSWK). In FFSWK, surface wave phase-delay sensitivity kernels  $K_{\alpha_V}^\phi$ ,  $K_{\alpha_H}^\phi$ ,  $K_{\beta_V}^\phi$ ,  $K_{\beta_H}^\phi$ ,  $K_\eta^\phi$  and  $K_\rho^\phi$  are calculated in a spherically symmetrical reference earth model, fully accounting for radial anisotropy. Surface wave phase delays can be written as (Zhou, 2009)

$$\delta\phi = \iiint_{\oplus} \left[ K_{\alpha_V}^\phi \delta \ln \alpha_V + K_{\alpha_H}^\phi \delta \ln \alpha_H + K_{\beta_V}^\phi \delta \ln \beta_V + K_{\beta_H}^\phi \delta \ln \beta_H + K_\eta^\phi \delta \ln \eta + K_\rho^\phi \delta \ln \rho \right] d\mathbf{x}^3, \quad (1)$$

where  $\alpha_H$  and  $\alpha_V$  are wavespeeds of horizontally-polarized and vertically-polarized P waves; and  $\beta_H$  and  $\beta_V$  are wavespeeds of horizontally-polarized and vertically-polarized shear waves; the parameter  $\eta$  is the speed of waves propagating in an intermediate direction and  $\rho$  is density.

In isotropic case, phase delays are can written as

$$\delta\phi = \iiint_{\oplus} \left[ K_\alpha^\phi \delta \ln \alpha + K_\beta^\phi \delta \ln \beta + K_\rho^\phi \delta \ln \rho \right] d\mathbf{x}^3, \quad (2)$$

where kernels for isotropic perturbations are simply the sum of kernels for anisotropic perturbations,

$$K_\alpha^\phi = K_{\alpha_V}^\phi + K_{\alpha_H}^\phi, \quad K_\beta^\phi = K_{\beta_V}^\phi + K_{\beta_H}^\phi. \quad (3)$$

The observable  $\delta\phi$  is phase difference between observed and reference seismograms. If we write the reference seismogram in the form  $s(\omega) = Ae^{-i\phi}$ , the observed displacement spectra can be written as  $o(\omega) = Ae^{-i(\phi+\delta\phi)}$ . In practice, we have to consider frequency averaging in spectra estimates which requires convolution of sensitivity kernels in frequency domain (Zhou *et al.* 2004, section 4.1; Zhou 2009, section 4). In FFSWK, we have used a fast computation scheme described in Zhou *et al.* (2004) to reduce frequency domain convolution to a simple time domain multiplication.

The following assumptions have been made: (1) synthetic and observed seismograms have been multiplied with a cosine taper centered at the group arrival of the reference fundamental-mode surface wave; and (2) the length of the cosine taper is at least several times longer than wave period. Fourier convention is the same as in Zhou *et al.* (2004), i.e.,

$$f(\omega) = \int_{-\infty}^{\infty} f(t)e^{-i\omega t} dt, \quad f(t) = \frac{1}{2\pi} \text{Re} \int_{-\infty}^{\infty} f(\omega)e^{i\omega t} d\omega. \quad (4)$$

## 2 Quick Start Guide

1. The programs have been successfully compiled with Intel FORTRAN Compiler. If you use a different compiler, there are likely I/O problems.
2. This is a first testing version and has not been benchmarked with any other code. **Use it at your own risk.**

The example below will show you how to quickly get started with FFSWK. Bold texts are Linux command lines.

1. Compile minos
  - **ifort minos\_bran.f -o minos\_bran**
2. run minos to generate eigenfunctions
  - **minos\_bran < minos\_bran.in.love**
  - **minos\_bran < minos\_bran.in.rayleigh**

The input file *minos\_bran.in.love* looks like this:

---

```
prem.aniso
love.out
eigf.love
1E-8 2
2
10 600 5 50 0 0
```

---

*prem.aniso*: file name of the reference earth model

*love.out*: output mode catalogue file (don't change the file name!)

*eigf.love*: unformatted binary output file of eigenfunctions (don't change the file name!).

*1E-8 2*: the first parameter controls integration precision and the second one defines lowest frequency in millihertz (mHz) above which gravitational terms are neglected.

*2*: toroidal modes only (for Love waves).

*10 600 5 50 0 0*: define modes to be calculated. (10 600): range of angular degree and (5 50): range of frequency in mHz. (0 0): fundamental mode  $n = 0$  only.

3. Compile and run program read\_eigf to prepare eigenfunctions for kernel calculation.
  - **ifort -132 -w -o read\_eigf read\_eigf.f**
  - **read\_eigf < read\_eigf.in.love**
  - **read\_eigf < read\_eigf.in.rayl**

The command line *read\_eigf < read\_eigf.in.love* reads toroidal mode eigenfunctions *eigf.love* and normalizes them. Normalized eigenfunctions are stored in directory *eigfs*.

The command line *read\_eigf < read\_eigf.in.rayl* reads spheroidal mode eigenfunctions *eigf.rayleigh* and normalizes them. Normalized eigenfunctions are stored in directory *eigfs*.

#### 4. Compile and run kernel calculation programs

- **ifort -132 -o ffswk ffswk.f90 ffswk-sub.f90**
- **ffswk < ffswk.input.love**

The input file ffswk.input.love looks like this:

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87	# mode 1 (~10 mhz) (see file love.out)
1	# love=1 rayleigh=2
eigfs	
182.58 -28.83 14.00	# source lon, lat, depth
7.130 -0.640 -6.500 0.860 3.790 -1.050	# moment Mrr Mtt Mpp Mrs Mrp Mtp
288.5090 42.611	# receiver lon, lat
600	# measurement window length in seconds
2 16	# Fresnel_number, sampling rate in Fresnel zone
200	# record sensitivity in the uppermost 200 km

---

when *Fresnel\_number*=1, only sensitivity in the first Fresnel zone will be calculated; when *Fresnel\_number*=2, the calculation domain is twice the size of the first Fresnel zone. *sampling rate in Fresnel zone* determines kernel spacial sampling rate. In general, 16–30 samples per Fresnel zone will be sufficient for most applications.

Output kernel files are:

kern/example.kernel.ah (this is $K_{\alpha_H}^\phi$ )	kern/example.kernel.av (this is $K_{\alpha_V}^\phi$ )
kern/example.kernel.bh (this is $K_{\beta_H}^\phi$ )	kern/example.kernel.bv (this is $K_{\beta_V}^\phi$ )
kern/example.kernel.et (this is $K_\eta^\phi$ )	kern/example.kernel.rh (this is $K_\rho^\phi$ )

The kernel files are data columns formatted as follow:

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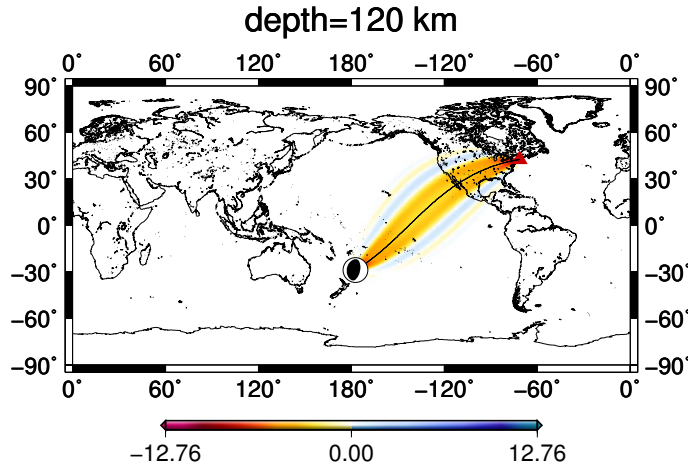
longitude	latitude	kernel_value	depth_in_km
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#### 5. Plot mapviews of sensitivity kernels

- **plot-kernel.pl kern/example.kernel.bh 120**

This command plots the kernel file *kern/example.kernel.bh* ( $K_{\beta_H}^\phi$ ) at a depth of 120 km. You should have a postscript file *kernel.xyz.ps* as shown in the Figure below. The unit of the sensitivity is  $10^{-8} \times \text{km}^{-3}$ .



### 3 Calculate your own sensitivity kernels

- To change earthquake source, seismic station or frequency of surface wave, simply go back to step 4 and modify the input file *ffswk.input.love* (or *ffswk.input.rayl*).
- To calculate kernels for a different earth model, go back to step 2 and replace the model file *prem.aniso* in *minos\_bran.in.love* (or *minos\_bran.in.rayleigh*) with your own model file.

### 4 References

1. Ying Zhou (2009), “Multi-mode 3-D surface-wave sensitivity kernels in radially anisotropic media”, *Geophysical Journal International*, **176**, 865-888.
2. Ying Zhou, F. A. Dahlen and Guust Nolet (2004), “3-D sensitivity kernels for surface-wave observables”, *Geophysical Journal International*, **158**, 142-168.
3. To calculate synthetic seismograms using Mineos, please refer to the CIG Mineos Webpage: <http://www.geodynamics.org/cig/software/mineos>
4. FFSWK Webpage: <http://seismo.geos.vt.edu/software.html>

### 5 Report Bugs!

This package does not come with any warranty or technical support. However, if you find bugs, please report them to Ying Zhou ([ying.yingzhou@gmail.com](mailto:ying.yingzhou@gmail.com)) and we will try to fix them and send updates to registered users.